Grid-free particle method for the inhomogeneous Enskog equation and its application to a Riemann-problem

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ABSTRACT. - Starting from the mollified version of the Enskog equation for a hard-sphere fluid, a grid-free algorithm to obtain the solution is proposed. The algorithm is based on the finite pointset method. For illustration, it is applied to a Riemann problem. The shock-wave solution is compared to the results of Frezzotti and Sgarra where good agreement is found. © Elsevier, Paris.

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1. Introduction

A kinetic equation describing the behaviour of moderately dense gases was first proposed by Enskog (1922) for the one-particle distribution function $f: \mathbb{R}_+ \times \Omega \times \mathbb{R}^3 \to \mathbb{R}_+$:

(EE)
$$\frac{\partial f}{\partial t} + v \cdot \nabla_x f = J_E(f)$$

and

(1)
$$J_{E}(f) = \iint_{\mathbb{R}^{3} \times S^{2}} k(v_{21} \cdot \eta) \Big[g^{(2)}(t, x, x + a\eta) f(t, x, v') f(t, x + a\eta, v'_{*}) - g^{(2)}(t, x, x - a\eta) f(t, x, v) f(t, x - a\eta, v_{*}) \Big] d\eta dv_{*}$$

In Eq. (1), a is the diameter of the molecules, $v_{21} = v_* - v$ is the relative speed of the centers, η is the unit vector from the center of the second sphere (*-indexed) to the center of the first, $k(v_{21} \cdot \eta) = a^2(v_{21} \cdot \eta)\theta(v_{21} \cdot \eta)$ for hard spheres, where θ is the Heaviside function. The primed variables denote the postcollisional values:

(2)
$$v' = v + (v_{21} \cdot \eta)\eta, \quad v'_* = v_* - (v_{21} \cdot \eta)\eta.$$

 $g^{(2)}$ is the pair-correlation function which accounts for spatial correlations. In the context of the revised Enskog theory (RET), $g^{(2)} = g^{(2)}(x, x_*|n(t))$ is the local equilibrium pair-correlation function expressed as a function of the number density n(t). For further detail refer to Résibois (1978).

In the framework of the so called Standard Enskog Theory (SET), $g^{(2)}$ is simply set equal to the value of the equilibrium pair-correlation function at contact, evaluated as a function of the uniform density n at the point of contact, where $g^{(2)} = g^{(2)}(n(x + \frac{a}{2}\eta))$.

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This work is concerned with the SET, where the equilibrium pair-correlation is such that the equation of state proposed by (Carnahan and Starling, 1969) is recovered:

(3)
$$g^{(2)}(\nu) = \frac{2-\nu}{2(1-\nu)^3}$$

where $\nu = \frac{bn}{4}$ is the reduced density and $b = \frac{2}{3}\pi a^3$. Approximation (3) agrees very well with the molecular dynamics data of Alder and Wainwright (1960), even in the metastable region. In contrast to the Boltzmann collision integral, the molecules now have a diameter a > 0 and thus the collision partners are at different positions. In the limit $\nu \to 0$, the pair-correlation $g^{(2)} \to 1$ and then for $a \to 0$, (1) recovers the Boltzmann collision integral for hard spheres.

Up to now, only a few attempts have been made to solve the Enskog equation numerically by kinetic schemes. Frezzotti and Sgarra (1993) combine a finite difference scheme with Monte Carlo techniques to evaluate the collision integral and apply it to a Riemann-problem for slightly dense gases. An algorithm which extends Bird's (1994) null-time-counter method to dense gases has been developed by Montanero and Santos (1997). They apply their Enskog simulation Monte Carlo (ESMC) method to a uniform shear flow.

Very recently, (Frezzotti, 1997) extended Bird's DSMC method to calculate the density profile of a dense gas in equilibrium near a hard wall.

In this paper a grid-free particle method is presented, which is an extension of Nanbu's scheme (Nanbu, 1980). Care is taken that every single collision conserves momentum and energy and they are thus globally conserved. It is based on the mathematically sound finite pointset method (FPM), which is sustained by the weak formulation of the discretised Enskog equation. The introduction of a mollified collision operator allows a rather generalized formulation for different implementations. The separation of space into a time-constant rigid grid is included as well as a local smoothing 'window' which is located around the second collision partner.

The organisation of the paper is as follows: At first, a measure formulation for the mollified time-discrete Enskog equation is derived in Section 2. In Section 3., measure approximation leads to an algorithm (particle method) for the solution of (EE). As an illustration, the method is applied to a Riemann problem in Section 4, where the formation of shock-profiles can be observed. The shock wave solutions are discussed and compared to Frezzotti's results. Finally, concluding remarks are given in Section 5.

2. Measure formulation for the time-discrete mollified Enskog equation

With the help of the one-particle distribution f(t, x, v) the macroscopic moments of f are obtained, e.g. the number density n, the mean velocity u or the temperature T

(4)
$$n(t,x) = \int f(t,x,v)dv$$

(5)
$$u(t,x) = (n(t,x))^{-1} \int v f(t,x,v) dv$$

(6)
$$T(t,x) = (3n(t,x)R)^{-1} \int |v - u(t,x)|^2 f(t,x,v) dv$$

where R is the gas constant.

Similar to the numerical approach to the full Boltzmann equation, a time splitting of (EE) into a free flow and a collision term is established:

For this, a time step $\Delta t > 0$ is chosen, $t_k = k \Delta t$, and $f^k = f(t_k, x, \tau)$, k = 0, 1, ... The splitting scheme is then

(7)
$$\frac{\partial \bar{f}^k}{\partial t} = -v \cdot \nabla_x f^k,$$

(8)
$$\frac{\partial f^k}{\partial t} = J_E(\bar{f}^k).$$

As it is well-known how to handle the free-flow part (7), we concentrate on the non-homogeneous collisional part (8).

Now let $\Omega \subset \mathbb{R}^3$ be the domain of position and [0,T] the time-domain. Then let

(9)
$$N_{\Omega}^{t} := \int_{\Omega} n(t, x) dx \quad < \quad \infty$$

be the number of molecules in Ω . The time-discrete formulation of (8) over a time-step Δt is then

(10)
$$f(t + \Delta t, x, v) = f(t, x, v) + \Delta t J_E(f(t, ., .))$$

Let $\varphi: \Omega \times \mathbb{R}^3 \to \mathbb{R}$ be a continuous test function. Multiplication of (10) and integration over the phase-space $\Omega \times \mathbb{R}^3$ leads to

(11)
$$\iint_{\Omega \times \mathbb{R}^{3}} \varphi(x,v) f(t+\Delta t,x,v) dv dx = \iint_{\Omega \times \mathbb{R}^{3}} \varphi(x,v) f(t,x,v) dv dx +$$

$$+ \Delta t \iiint_{\Omega \times \mathbb{R}^{3} \times \mathbb{R}^{3} \times S^{2}} k(v_{21} \cdot \eta) g^{(2)} \left(n \left(t, x - \frac{a}{2} \eta \right) \right) \left[\varphi(x,v') - \varphi(x,v) \right] \times$$

$$\times f(t,x,v) f(t,x-a\eta,v_{*}) d\eta dv_{*} dv dx$$

A usual approach to (11) in the rarefied gas regime is the projection of the initial density f(t,.,.) onto locally constant functions on an arbitrary partition of Ω and the application of a particle method working independently on each of these cells (homogenisation). In the dense gas case, the cells cannot be regarded as independent of each other anymore. Particles from different cells might interact, i.e. the whole particle ensemble has to be considered in the collision term. From this point of view, a parallelization of the collision term, as is easily done in the (cellwise independent) rarefied gas regime, is not straightforward. Yet, as the whole particle ensemble always has to be considered in the dense gas case, one might think of other grid structures than rigid grids. The approach, which includes a whole class of different structures (including the rigid grid), is shown in the following.

A suitable approach to the evaluation of the distribution function at different positions is found by the discretisation of the collision operator in the spatial domain with the help of a mollifier. The idea of smoothing the Boltzmann equation (BE) has been introduced by Morgenstern (1955) and extended by Steiner (1995) who derived numerical schemes for (BE) using mollifiers. This formulation is now applied to the Enskog equation, which changes the collision operator $J_E(f)$ in (1) to

(12)
$$J_{E}^{\beta}(f) = \iiint_{\mathbb{R}^{3} \times \mathbb{R}^{3} \times S^{2}} k(v_{21} \cdot \eta) \beta_{h}(x, x_{*}) \\ \times \left[g^{(2)} \left(n \left(t, \frac{1}{2} (x + x_{*} + a \eta) \right) \right) f(t, x, v') f(t, x_{*} + a \eta, v'_{*}) \right. \\ \left. - g^{(2)} \left(n \left(t, \frac{1}{2} (x + x_{*} - a \eta) \right) \right) f(t, x, v) f(t, x_{*} - a \eta, v_{*}) \right] d \eta d v_{*} d x_{*}$$

where $\beta_h(x,.)$ is, for any x, a mollifier with compact support $\Omega_{\beta}(x)$. The mollifier has to fulfill certain conditions such as symmetry, normalization and boundedness so that $J_E^{\beta} \to J_E$ if $\beta_h(x,.)$ converges to Dirac's δ -function.

A change of variable $\tilde{x}_* = x_* - a\eta$ and the introduction of measures $d\mu^t = f(t, x, v)dvdx$ and $d\mu_*^t = f(t, \tilde{x}_*, v_*)dv_*d\tilde{x}_*$ delivers the time-discretised measure-valued Enskog equation (without free flow term):

(13)
$$\iint_{\Omega \times \mathbb{R}^{3}} \varphi(x,v) d\mu^{t+\Delta t} = \int_{S^{2} \times \Omega \times \mathbb{R}^{3} \times \Omega_{\beta}^{\eta}(x) \times \mathbb{R}^{3} \times [0,1]} \frac{1}{4\pi N_{\Omega_{\beta}^{\eta}(x)}^{t}} \left(\varphi(x,v') \mathcal{X}_{[0,\Delta t \ \tilde{q}_{21}]}(\kappa) + \varphi(x,v) \mathcal{X}_{[\Delta t \ \tilde{q}_{21},1]}(\kappa) \right) d\kappa d\mu_{*}^{t} d\mu^{t} d\eta.$$

The idea of representing the weak formulation of the collision integral via a characteristic function $\mathcal{X}_{[a,b]}(x) = 1$ if $x \in [a,b]$ and 0 otherwise, is due to Struckmeier (1994), (see also Neunzert and Struckmeier, 1997).

In equation (13), $\Omega^{\eta}_{\beta}(x) = \{y|y+a\eta \in \Omega_{\beta}(x)\}$ is the domain in which possible collision partners for a particle at position x are located, $N^t_{\Omega^{\eta}_{\beta}(x)}$ is the number of particles in this domain at time t, according to Definition (9), and

(14)
$$\tilde{q}_{21} = 4\pi N_{\Omega_{\beta}^{\eta}(x)}^{t} k(v_{21} \cdot \eta) \beta_{h}(x, \tilde{x}_{*} + a\eta) g^{(2)}(n(t, (x + \tilde{x}_{*})/2))$$

is the collision probability for two particles.

Equation (13) renders the measure of φ with respect to the density f at time $t + \Delta t$ as a function of the measure φ at time t with respect to a product of density functions, i.e. if the measures μ and μ_* are known at time t, then the measure μ at time $t + \Delta t$ is computable via (13).

Equation (13) is reasonable in the sense of measures, i.e. non-negativity of $f(t + \Delta t)$ is guaranteed for non-negative f(t), if

(15)
$$\Delta t \ \tilde{q}_{21} \le 1 \qquad \forall \text{ admissible } (x, \tilde{x}_*, \eta, v_{21} \cdot \eta).$$

This is a restriction for the time discretisation Δt in the Euler scheme!

Remark 2.1.

1. If $\bigcup_{k \in \mathbb{K}}^{\circ} Z_k^h = \Omega$ is a disjoint partition of Ω with diameter $\operatorname{diam}(Z_k^h) = \sup\{\|x - x_*\|, x, x_* \in Z_k^h\} \le h$ for all $k \in \mathbb{K}$ then

(16)
$$\beta_h^{hom}(x, x_* + a\eta) = \frac{\sum_k \mathcal{X}_{Z_k^h}(x) \mathcal{X}_{Z_{k^*}^h}(x_*)}{\sum_k \mathcal{X}_{Z_k^h}(x) \cdot \text{meas}(Z_{k^*}^h)}$$

is a grid-inducing mollifier allowing collisions of particles of different cells. $Z_{k^*}^h$ is the cell which contains the position $x-a\eta$. A mollifier of this kind where Z_k^h are cubic cells of (constant) length h has been used by Montanero $et\ al$. Typically, h is smaller than the mean free path.

2. In the implementation based upon the above formulation, a local mollifier is used,

(17)
$$\beta_h^{loc}(x, x_* + a\eta) = \beta_h^{loc}(y) = \frac{1}{h} \mathcal{X}_{[0, h/2]}(|y|), \quad y = x - x_* - a\eta,$$

where the mean free path $\lambda = \left(\sqrt{2}\pi na^3g^{(2)}(n)\right)^{-1}$ is taken as the maximum value for h: $h \leq \lambda$. (The form of the mean free path here is rather conventional, since it is the value of the equilibrium mean free path.) The mollifier β_h^{loc} allows a particle at position x to find a collision partner in an environment of $x-a\eta$ with no regard to a grid. Essentially, this mollifier is a self-adaptive 'window' of size h, where h is directly coupled to the local mean free path. A large h may result in greater effort for the bookkeeping of the particles (the higher the number of particles in the window, the greater the cost of choosing a collision partner randomly) while with too small an h, one might not find a collision partner at all. In the following examples care was taken, that at the beginning of the collision process, a minimum of 200 particles were located within any window.

In contrast to the rigid grid above, the local mollifier already includes a flexible, time- and location-dependent treatment of the particles. On the other hand, the computational costs might be higher, due to a more sophisticated bookkeeping of the particles. The efficiency of rigid grids and locally acting mollifiers have, to the knowledge of the author, not yet been compared in the dense gas regime.

3. Particle method for the Enskog equation

In this section the principles of particle approximations are introduced and applied to (13). This delivers an algorithm for the numerical solution of one time step.

3.1. Particle approximation

A particle is characterized by its position x, velocity v and a weight 1/N, where N is the number of particles. A particle set (or finite point set) is given by

(18)
$$\mu_N = \{(1/N, x_1, v_1), \dots, (1/N, x_N, v_N)\}\$$

or - in different notation - by

(19)
$$\delta_{\mu_N} = \frac{1}{N} \sum_{i=1}^N \delta_{x_i} \delta_{v_i}$$

Now, for a given density $f \in \mathcal{L}^1_+$ we say that ' δ_{μ_N} converges to f', if

(20)
$$\lim_{N \to \infty} \frac{1}{N} \sum_{i=1}^{N} \varphi(x_i, v_i) = \int f \cdot \varphi \, dv dx, \quad \text{for all } \varphi \in C^b(\Omega \times \mathbb{R}^3).$$

This means that the discrete measure δ_{μ_N} converges weakly* (i.e. convergence in the dualspace of $C^b(\Omega \times \mathbb{R}^3)$) to fdvdx. Eq. (20) can be interpreted as an integration rule, where we integrate the function φ with respect to the measure fdvdx. The idea of particle methods is now to approximate continuous measures by discrete ones with the help of (20). Approximating the right-hand side of eq. (13) with respect to some distance between μ_N and f (discrepancy) thus leads to a finite pointset method for the approximation of the distribution function f at time f at time f and f (discrepancy) thus leads to a finite pointset method for the approximation of the distribution function

Remark 3.1. – Further detail to the approximation of continuous measures by finite point sets and the quality of approximation can be found in Struckmeier (1994). He also states how to include unbounded φ such as $|v|^2$ or $|v|^2v$ etc.

3.2. FINITE POINTSET METHOD FOR THE INHOMOGENEOUS ENSKOG EQUATION

Concerning the discretisation of the inhomogeneous equation we use a decoupling of the free flow of particles and the collisions among them (splitting scheme (7) & (8)).

Free Flow: Given an approximation of $f(k\Delta t, x, v)$ by a finite pointset we use (7) to obtain

(21)
$$\bar{f}((k+1)\Delta t, x, v) = f(k\Delta t, x - \Delta t v, v),$$

i.e. we just move the particles over the time increment Δt with the particle velocity and no spatial discretisation is required.

Collision Term: Let $\mu_N = \{(1/N, x_1, v_1), \dots, (1/N, x_N, v_N)\}$, $x_i \in \Omega$, $v_i \in \mathbb{R}^3$ be a particle approximation of $\bar{f}dvdx$. Then we construct an approximation of the product measure $\kappa \otimes \mu_* \otimes \mu \otimes \eta$ and evaluate (13) for the discrete product measure. The simulation scheme is then:

- 1. For any particle x_i choose a collision parameter η_i uniformly distributed in S^2 .
- 2. Choose uniformly a test collision partner x_{i^*} which is situated in the support of $\beta_h(x_i, +a\eta_i)$.
- 3. Choose a uniformly distributed $\kappa_i \in [0, 1]$.
- 4. Determine the post-collisional state (2) of (x_i, v_i) and (x_{i^*}, v_{i^*}) at time $(k+1)\Delta t$:

(22)
$$(v_i^{k+1}, v_{i^*}^{k+1}) = \begin{cases} ((v_i^k)', (v_{i^*}^k)') & \text{if } \kappa_i \leq \Delta t \ \tilde{q}_{21} \\ (v_i^k, v_{i^*}^k) & \text{otherwise.} \end{cases}$$

5. Repeat until no more test collision pairs can be found.

Applying the two steps above for each time step Δt , a solution scheme for the inhomogeneous and instationary Enskog equation arises.

Remark 3.2.

- 1. For the rarefied gas case, where the Enskog equation reduces to the Boltzmann equation, a convergence proof for Nanbu's method can be found in Babovsky and Illner (1989).
- 2. As every single collision conserves momentum and energy, these moments are globally conserved in the scheme.

In the following section, this algorithm is applied to the one-dimensional Riemann problem, where shock-wave solutions appear.

4. Application of the scheme to the 1d-Riemann problem

In the present section the propagation of a plane shock wave in a dense gas is studied. Attention is focused on the calculation of the fully formed shock profile, which appears as stationary to an observer moving with

the shock front. The x-direction corresponds to the e^1 -direction in cartesian coordinates. It is assumed that far from the shock equilibrium conditions exist, i.e.

(23)
$$\lim_{x \to -\infty} f(x, v) = \frac{n_1}{(2\pi RT_1)^{3/2}} \exp\left(-\frac{(v - u_1 e^1)^2}{2RT_1}\right)$$

(24)
$$\lim_{x \to +\infty} f(x, v) = \frac{n_2}{(2\pi R T_2)^{3/2}} \exp\left(-\frac{(v - u_2 e^1)^2}{2R T_2}\right)$$

and $x, n_i, T_i, u_i \in \mathbb{R}^1$ and $v, e^1 \in \mathbb{R}^3$. The parameters of the downstream and upstream equilibrium states are connected by the following Rankine-Hugoniot relationships:

(25)
$$n_1 u_1 = n_2 u_2$$

$$n_1 [u_1^2 + RT_1(1 + bn_1 g^{(2)}(n_1))] = n_2 [u_2^2 + RT_2(1 + bn_2 g^{(2)}(n_2))]$$

$$u_1^2 + RT_1(5 + 2bn_1 g^{(2)}(n_1)) = u_2^2 + RT_2(5 + 2bn_2 g^{(2)}(n_2))$$

which arise from the equality of upstream and downstream fluxes of mass, momentum and energy. Due to the influence of the correlation function $g^{(2)}$ in (25), the ratio n_1/n_2 , for constant upstream bulk velocity u_1 , now varies with the reduced density ν , which is not the case for a rarefied gas. The upstream mean velocity is normalized: $M = u_1/(\gamma RT_1)^{1/2}$ ($\gamma = 5/3$). It should be noted that M is not the effective Mach number, because the adiabatic speed of sound depends on ν , too.

4.1. Description of the numerical environment

Each computation started with the gas filling the half-space x < 0 in the upstream equilibrium condition and the corresponding downstream equilibrium condition in the half-space x > 0. At the end of the instationary part of the system's evolution, an averaging over many time-steps took place to smoothen the oscillations of the discrete measures.

Computational parameters were 1000 initial particles per mean free path λ , the support of the mollifier was $h=0.2\lambda$ and the dynamics were evaluated for 1000 time steps, where one time step is equal to $\lambda/\sqrt{2RT_1}$, the mean free time between collisions. From the FPM-numerics for the Boltzmann equation it is already well known that the (numerical) shock does not move at constant speed, see e.g. Giering (1995). Instead, the shock wave oscillates around a mean point. This means that the solution has to be averaged over some 50-60 runs, or the shock wave has to be traced during the dynamics. In the present implementation, a tracing of the shock wave took place, which delivered quite good results, but for the low hypersonic case. For a Mach number M=2, the resulting shock gradients in the rarefied regime and for $\nu=0.04$ are 5 % and 10 % larger than the prediction of Mott-Smith's v^2 -curve and Frezzotti's results, respectively.

4.2. Numerical results

A first series of calculations considered the variation of the shock-profile for a constant upstream velocity but varying density. The result is presented in Figures 1-3, showing the density, temperature and velocity plots of the shock-wave solution. To help comparing shock profiles in different conditions, each quantity is normalized to its own variation across the shock region. As is easily seen, the width of the shock grows with increasing ν . Even with scaling to a constant length $\lambda_B = (\sqrt{2}\pi na^3)^{-1}$, which is not pictured here, the widening of

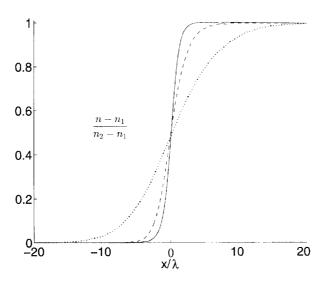


Fig. 1. – Normalized density profile for M=4 and varying ν : Solid line, $\nu=10^{-10}$. Dashed line, $\nu=0.1$. Dotted line, $\nu=0.2$.

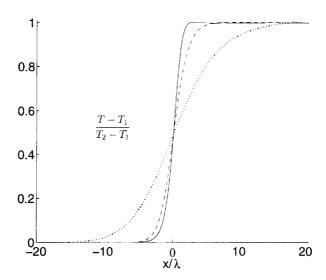


Fig. 2. – Normalized temperature profile for M=4 and varying ν : Solid line, $\nu=10^{-10}$. Dashed line, $\nu=0.1$. Dotted line, $\nu=0.2$.

the shock is noticable. As in the rarefied gas case, the temperature profile precedes the density profile, see Figure 4 for M=4 and $\nu=0.1$.

A second interest was taken in the dependency of the shock-profile on the upstream velocity if ν was fixed. From the theory of rarefied gases it is well-known that for small shock-speeds the maximum gradient of the profiles increases with the shock-speed, see e.g. Mott-Smith (1951). The same phenomenon appears to be true in the dense gas limit as well. In Figure 5, the reduced density was taken as $\nu = 0.1$ and the shock-velocity varied.

Further calculations have been made to allow a comparison to already known results. For this the concept of the reciprocal shock thickness δ^{-1} is used, which is defined as the maximum value of the normalized

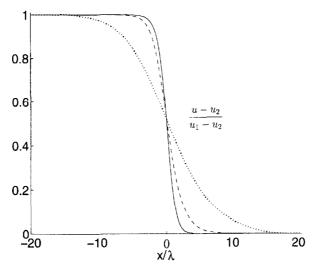


Fig. 3. – Normalized velocity profile for M=4 and varying ν : Solid line, $\nu=10^{-10}$. Dashed line, $\nu=0.1$. Dotted line, $\nu=0.2$.

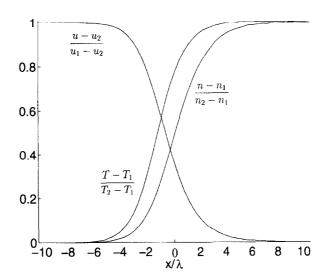


Fig. 4. – Normalized density, temperature and velocity profiles for M=4 and $\nu=0.1$: The density shock is preceded by the temperature.

density gradient:

(26)
$$\delta^{-1} = \frac{1}{n_2 - n_1} \max\left(\frac{dn}{dx}\right)$$

The data displayed in Figure 6 have also been obtained by a time average of many time steps after the initial discontinuity had changed into a stationary profile. From the averaged profile the maximum density gradient was extracted. In the rarefied limit $\nu \to 0$, the values predicted by Mott-Smith (1951) were recovered quite well except for M=2, where the shock oscillated within a region of length 1.6λ (in contrast to a fluctuation region of length 0.8λ for M=4). The high fluctuations in the low hypersonic regime might be the reason for the remarkable deviation of the shock thickness from theory.

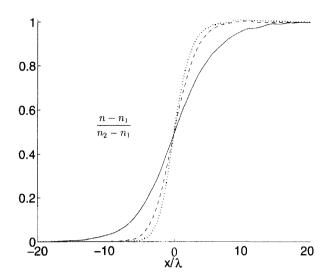


Fig. 5. – Normalized density profile for $\nu=0.1$ and varying M: Solid line, M=2. Dashed line, M=3. Dotted line, M=4.

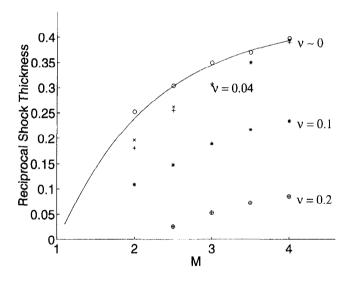


Fig. 6. – Reciprocal shock thickness as a function of M and ν . The solid line represents Mott-Smith v^2 curve for a hard-sphere potential. Symbols: \bigcirc , $\nu=10^{-10}$; \times , $\nu=0.04$ using Frezzotti's $g^{(2)}$; +, Frezzottis results; *, $\nu=0.1$; \oplus , $\nu=0.2$.

In a further series we checked with Frezzotti's results obtained from a different pair correlation function, see (Frezzotti and Sgarra, 1993), where in his notation E=0.4 was taken. This corresponds to a slightly dense gas with $\nu=0.04$. Also here the comparison is rather good for any calculated state except for low M.

For small ν , the data is quite close to the rarefied limit, which should be expected. It also seems to converge to the low density limit with increasing shock speed, at least for the tested value $\nu=0.04$. For larger ν , this becomes more and more untrue, the shock profile widens and thus the reciprocal shock thickness decreases.

5. Concluding remarks

In this paper a finite pointset method is presented which is able to handle the inhomogeneous Enskog equation for dense gases as well as for rarefied gases. The algorithm is based on the well-known finite pointset method and conserves energy and momentum. Further, it is open to either a grid-based scheme or a grid-less one, where for the examples a grid-less mollifier was taken. It is able to calculate the profile of shock waves even if the mean free path λ is smaller than the diameter of the molecules, e.g. for $\nu=0.2$ the diameter was $a=2.98\lambda$. The comparison to Frezzotti's results showed a rather good agreement.

The standard Enskog theory was used hereby to model the pair correlation function $g^{(2)}$. Here it would be interesting to observe the influence of replacing $g^{(2)}$ by the functional expression as derived by the revised Enskog theory (Résibois, 1978), which requires a higher computational effort. Also, a thorough comparison of rigid grids and local self-adaptive mollifiers is of computational interest.

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